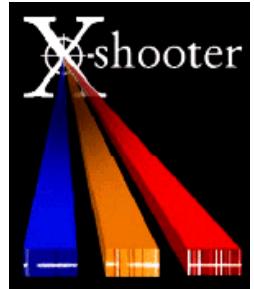




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A new arrival at the VLT: the commissioning of the X-shooter spectrograph

P. Di Marcantonio – INAF-OATs (on behalf of the X-shooter consortium)



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X-shooter consortium



X-shooter has been built by a Consortium of European Institutes comprising:

- ESO (*PI S. D'Odorico, PM H. Dekker*)
- Denmark (*PI/PM P. Kjærgaard-Rasmussen*)
- France (*PI F. Hammer, PM I. Guinouard*)
- The Netherlands (*PI L. Kaper, PM R. Navarro*)
- Italy (***PI R. Pallavicini, PM F. Zerbi***)

ATG has been responsible for the design, development, integration and commissioning for the control software of the whole instrument.

People involved @ INAF:

Roberto Pallavicini (Co-PI),

F. M. Zerbi, V. de Caprio, A. de Ugarte Postigo, M. Riva, P. Spanò, M. Tintori (INAF-OABr)

P. Di Marcantonio, P. Santin, A. Zacchei, C. Zamberlan, M. Vidali (INAF-OATs)

R. Cosentino, P. Bruno (INAF-OACt)



X-shooter resources



	effort (FTE)	Cost (k€)	Contribution
Denmark	19	850	Control electronics, Backbone unit, FEA, system test of UVB spectrograph
ESO	15	1510	Overall ProjMan.& SysEng, detectors, final system integration and commissioning, logistics
France	12	140	IFU, DRS
Italy	19	800	Optomechanical design and integration of UVB and VIS, system test of VIS spectrograph, Control Software
Netherlands	1.8	2044	NIR spectrograph, contribution to DRS
Total	66.8	5344	



X-shooter GTO



INAF-OATs (only) GTO involvement:

	Guaranteed Nights
Italy	44.5
Denmark	45.5
France	20.8
Holland	43.5
Total	154.3

- Abundances and Dust in high redshift ($z>4.0$) Damped Lyman α galaxies (PI: Molaro P., CoI: Vladilo G., D'Odorico V.)
- Optical-NIR spectra of quasars close to re-ionization ($z\sim 6$) (PI: D'Odorico V., CoI: Molaro P., Cristiani S., Viel M., Vladilo G.)
- Extremely metal-poor stars in SDSS fields (PI: Bonifacio P., CoI: Molaro P.)
- Tomography of the Intergalactic Medium with multiple QSO lines of sight (PI: Stefano C., CoI: F. Calura, E. Vanzella, V. D'Odorico, M. Viel, P. Monaco, ...)
- A 100 burst X-Shooter / Swift GRB afterglow legacy survey (CoI: Pian.E.)
- X-Shooting Supernovae (CoI: Pian E.)
- Study in situ of GRB progenitors and their host galaxies with X-Shooter : from $z = 0.1$ to 2.3 (CoI: Pian E.)



X-shooter time schedule



STC 56	Dec 2003
PDR	Dec 2004
FDR	Feb – Jul 2006
Integration @ ESO	2008
PAE	Sept 2008
Comm #1	Nov 2008
Comm #2	Jan 2009
Comm #3	Mar 2009
Comm #4	May 2009
SV or GTO (several periods)	Jul – Sept 2009
<u>Start of Operations</u>	1.10.2009



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X-shooter commissioning team



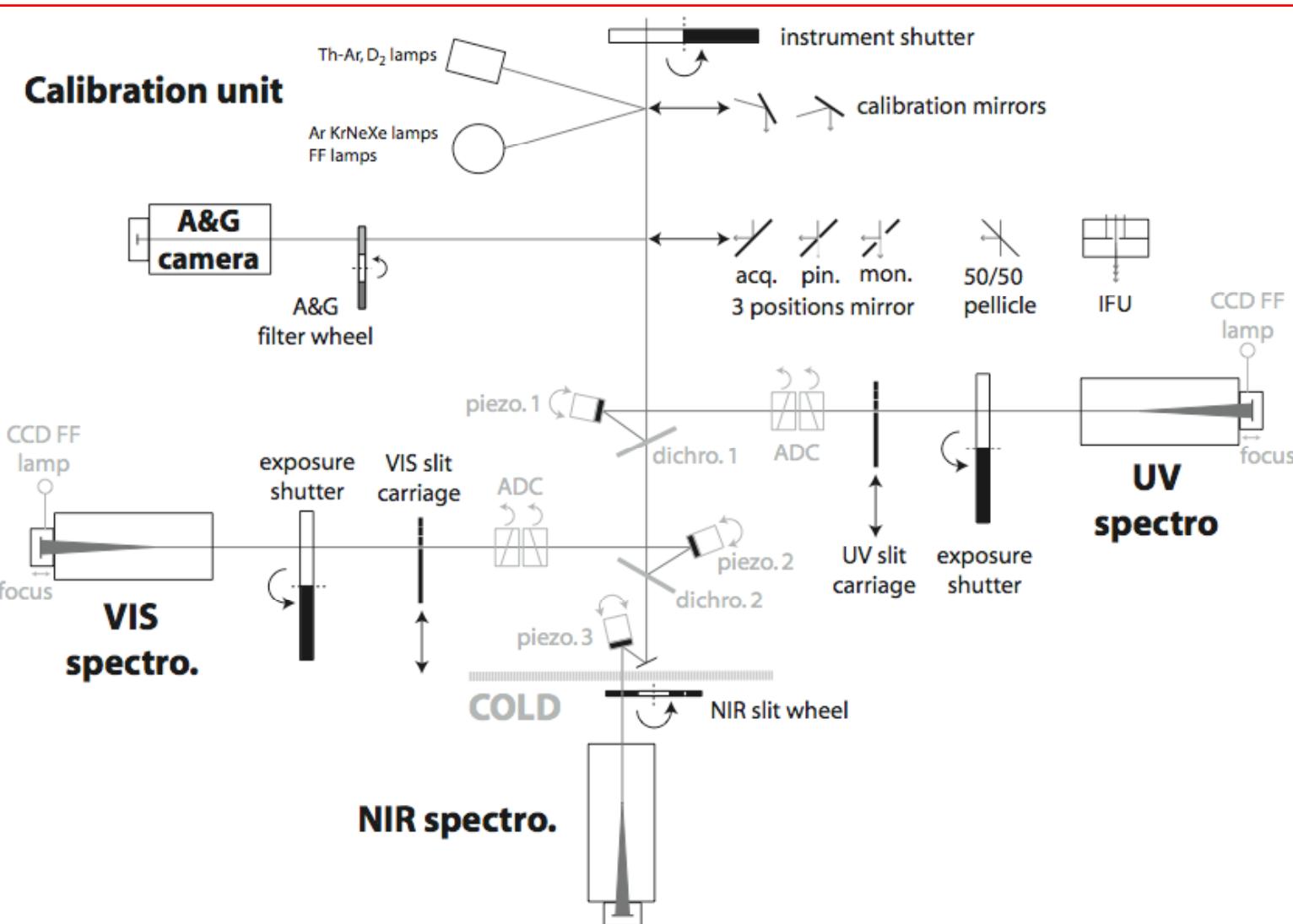


Instrument characteristics

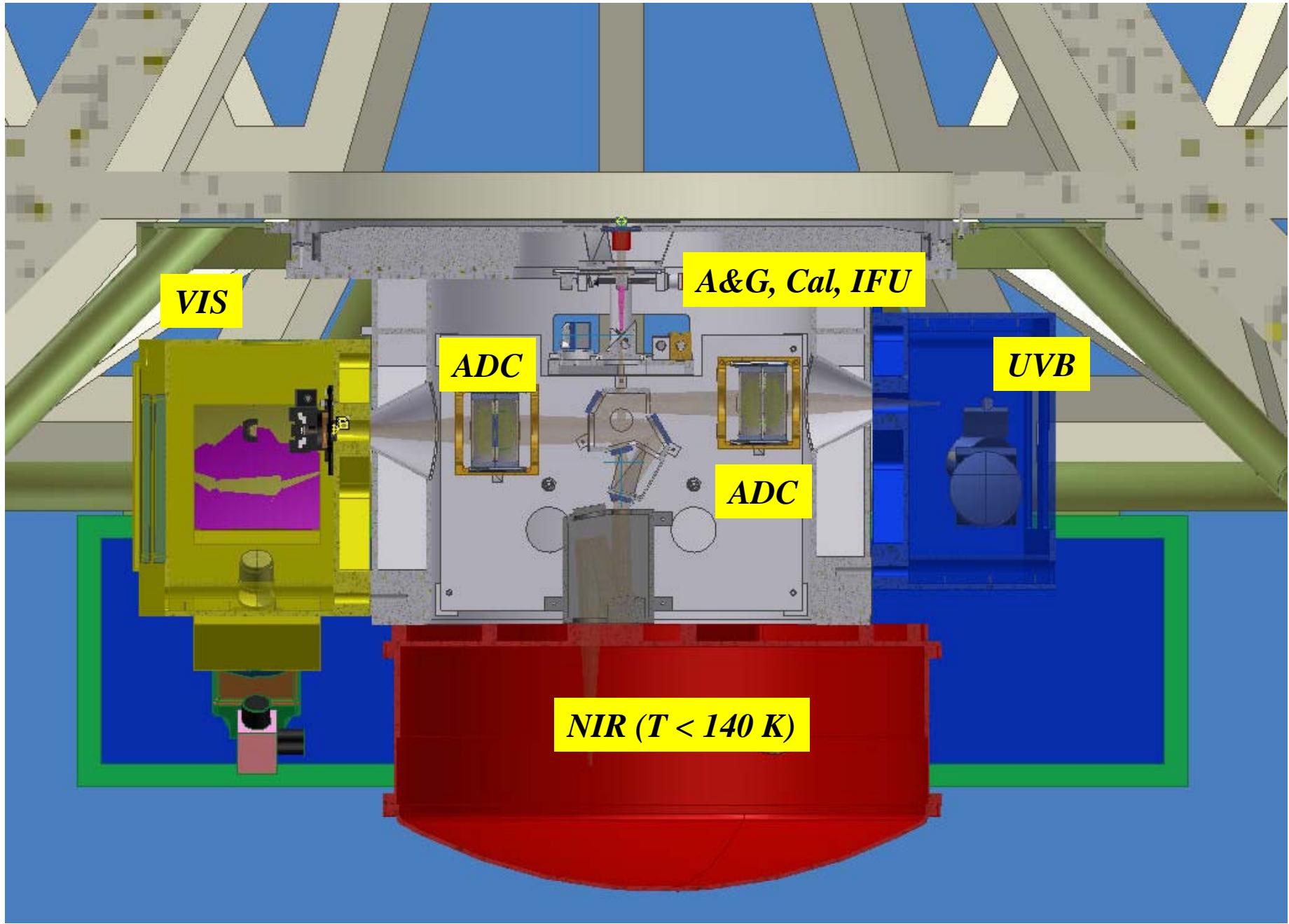


- Wavelength range: three arms covering from 300 nm to 2500 nm
- Fixed prism cross-dispersed echelle format (slit length 11'')
- Detectors:
 - 2K x 4K 15μ CCDs (UVB and VIS arms)
 - 2K x 1K segment of a 2K x 2K 18μ Hawaii 2RG MBE (NIR arm)
- IFU (1.8'' x 4''), ADC for UVB and VIS arms, calibration unit and A&G unit
- Spectral resolution: ~7000 to 12000 for 0.6'' slit or IFU
- High Detective Quantum Efficiency
- Pipeline delivering sky-subtracted, wav cal 2D spectra and 3D data cube for the IFU

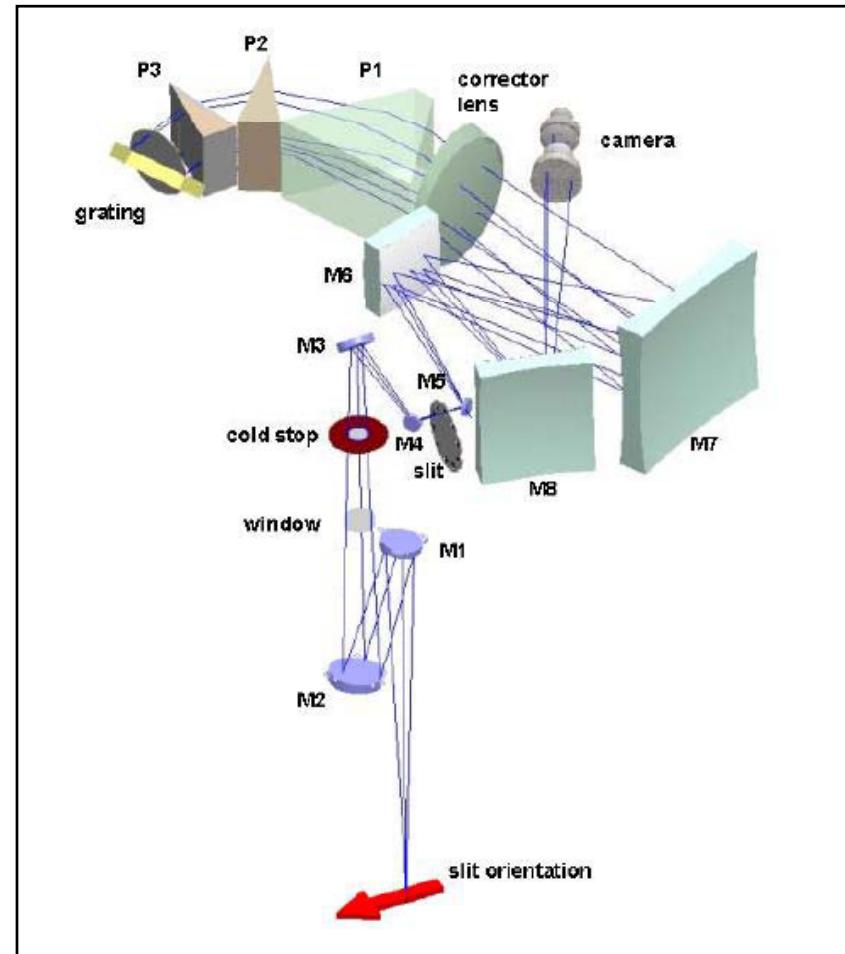
Instrument layout I

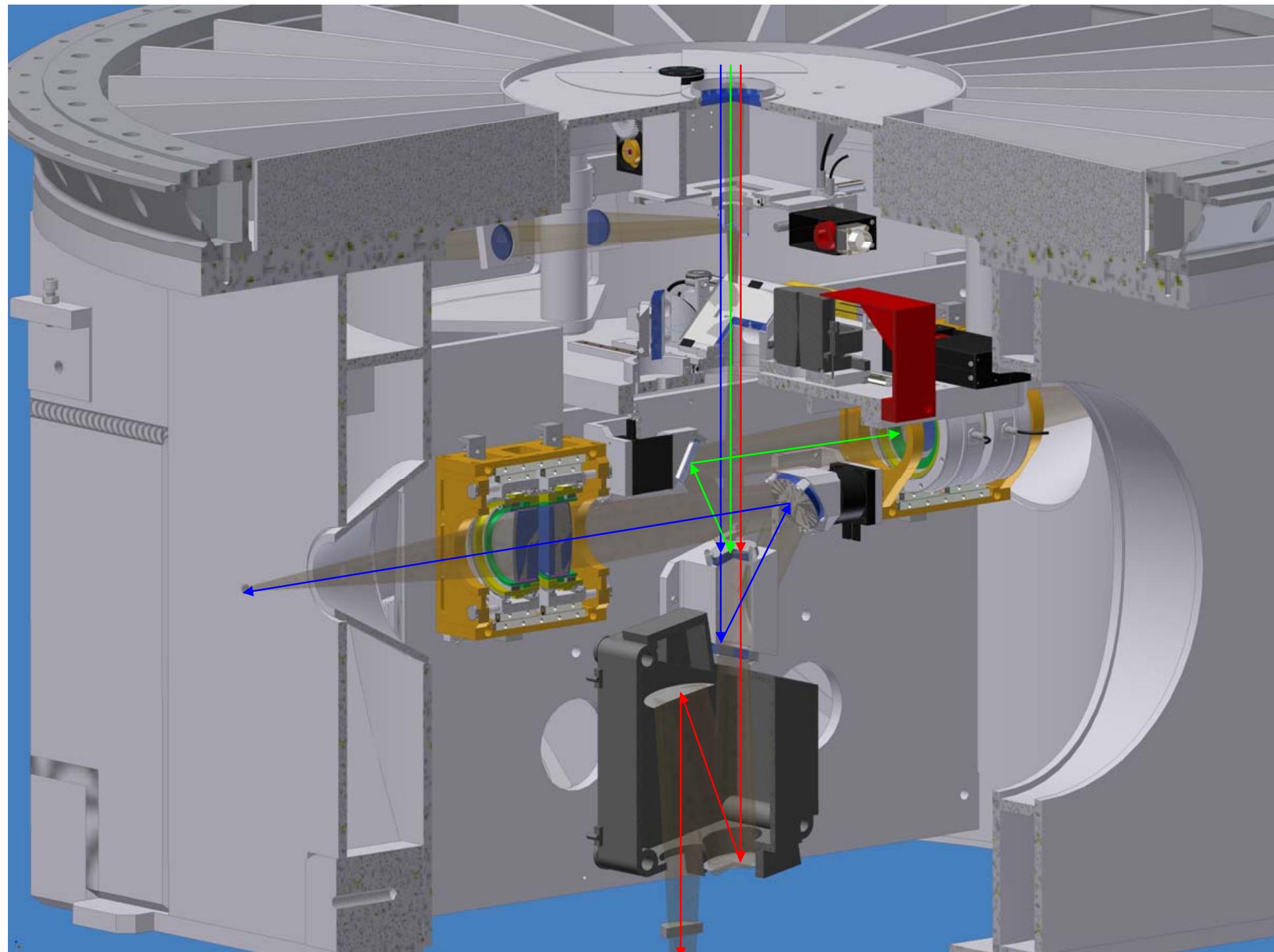


Instrument layout II

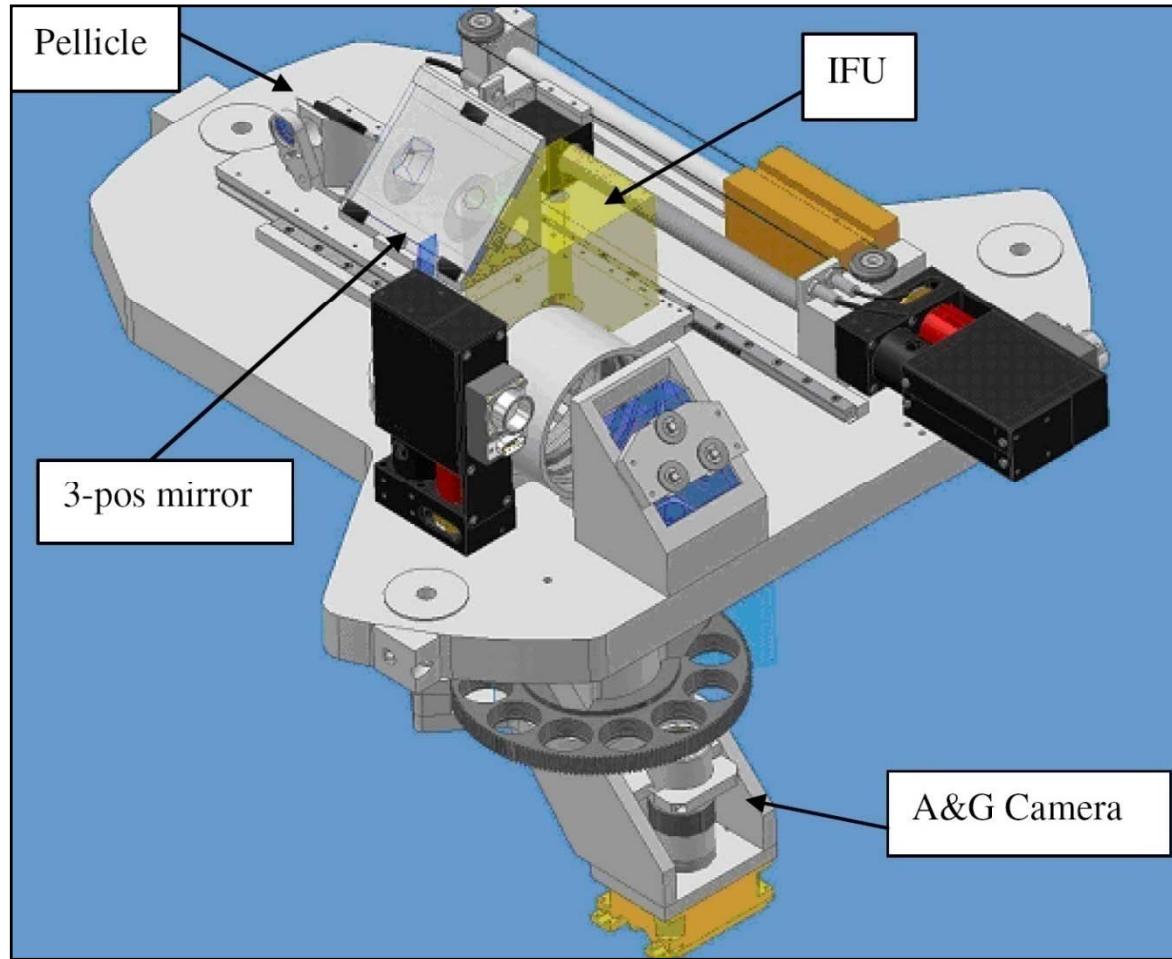


NIR optical layout

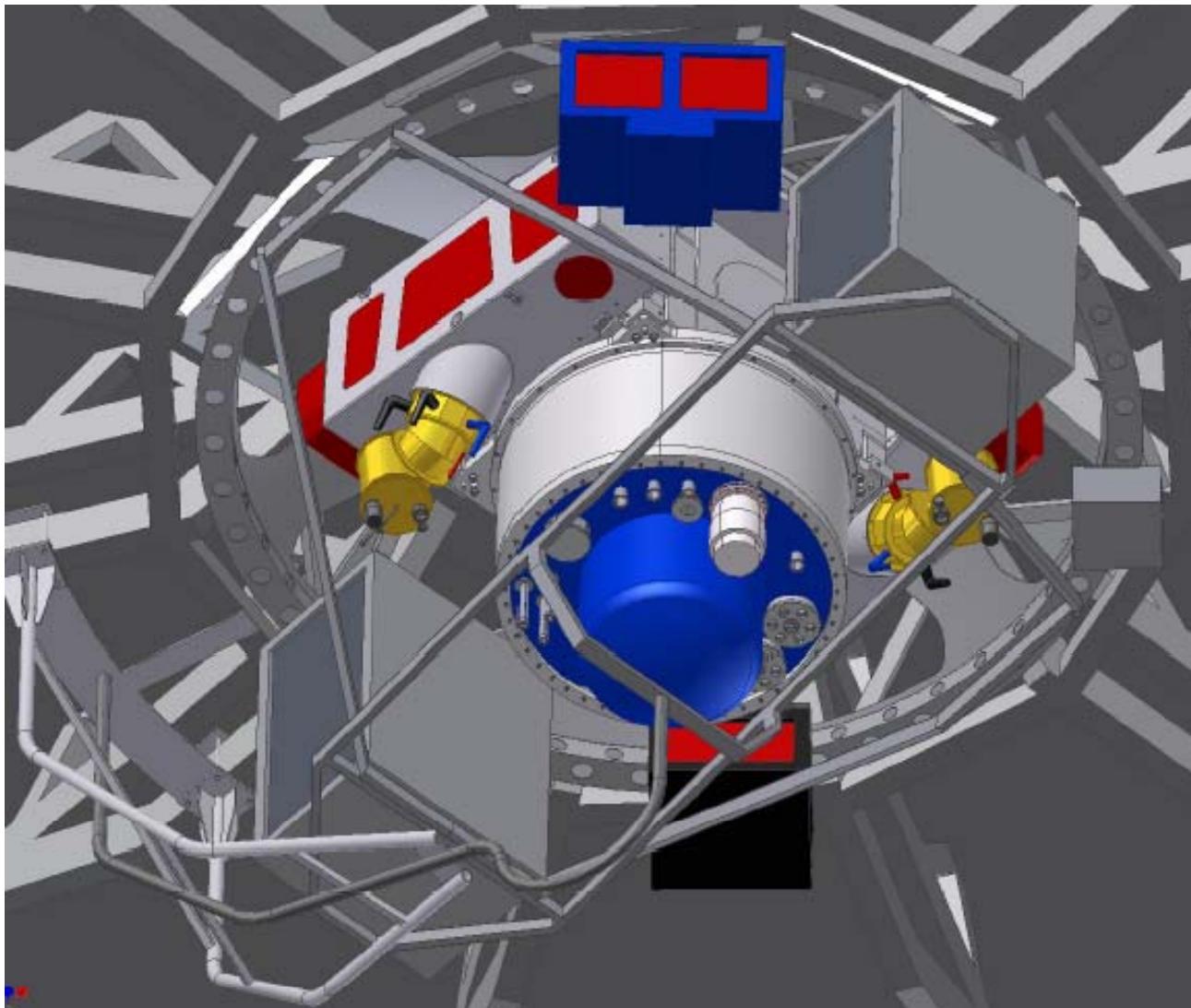




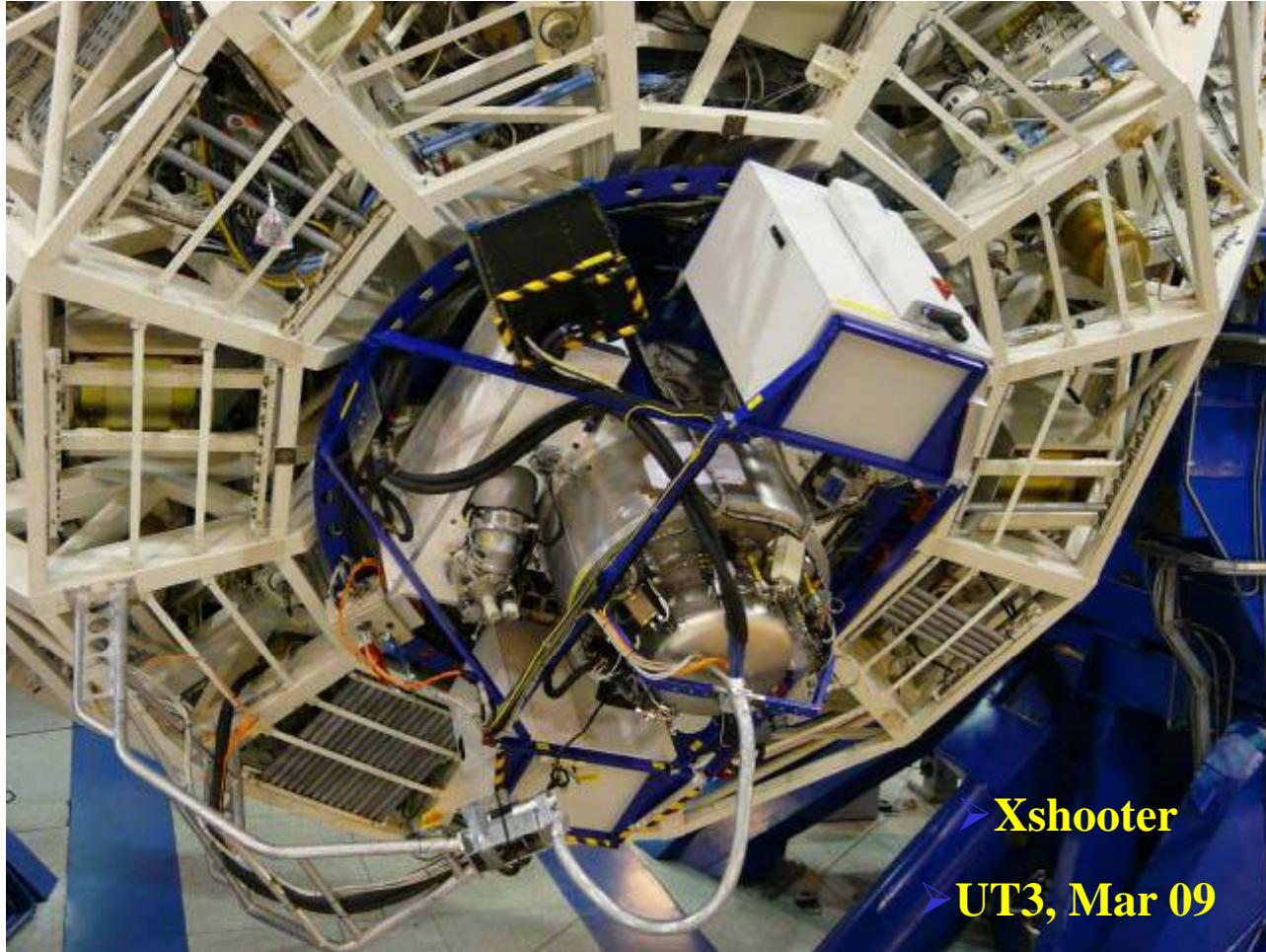
A&G unit



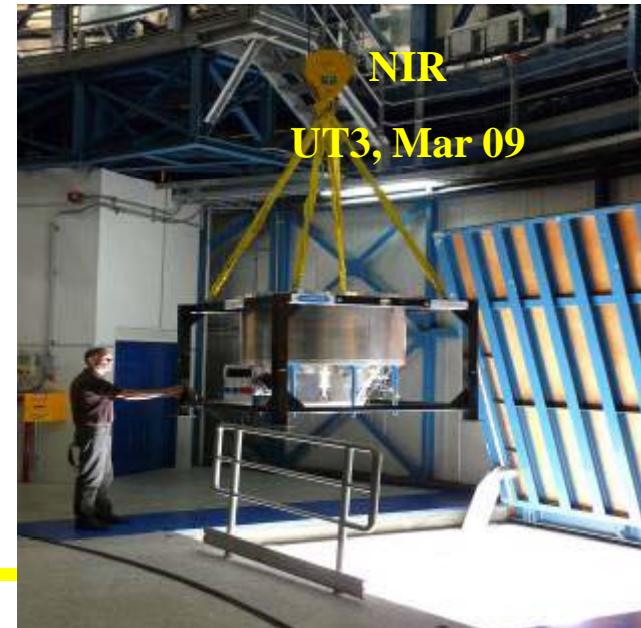
FDR CAD design

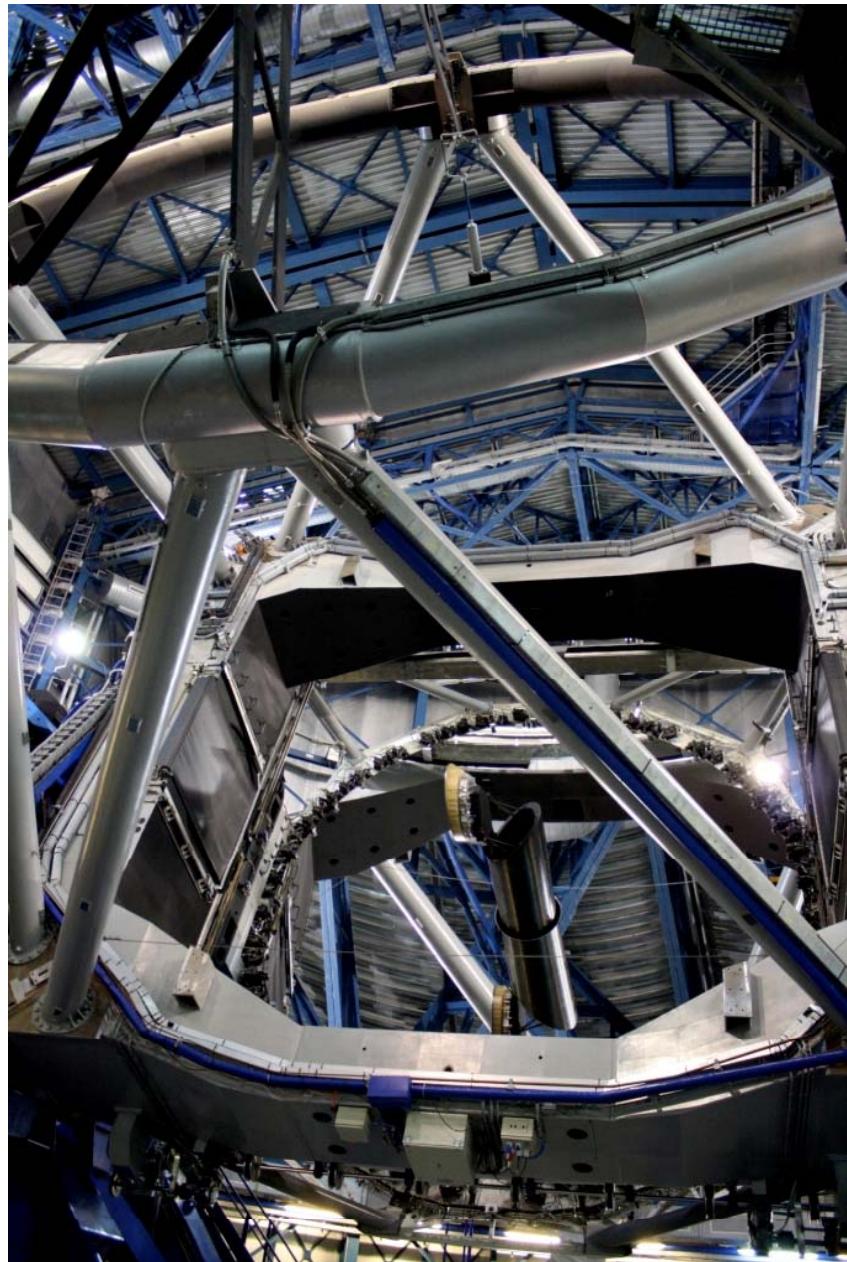
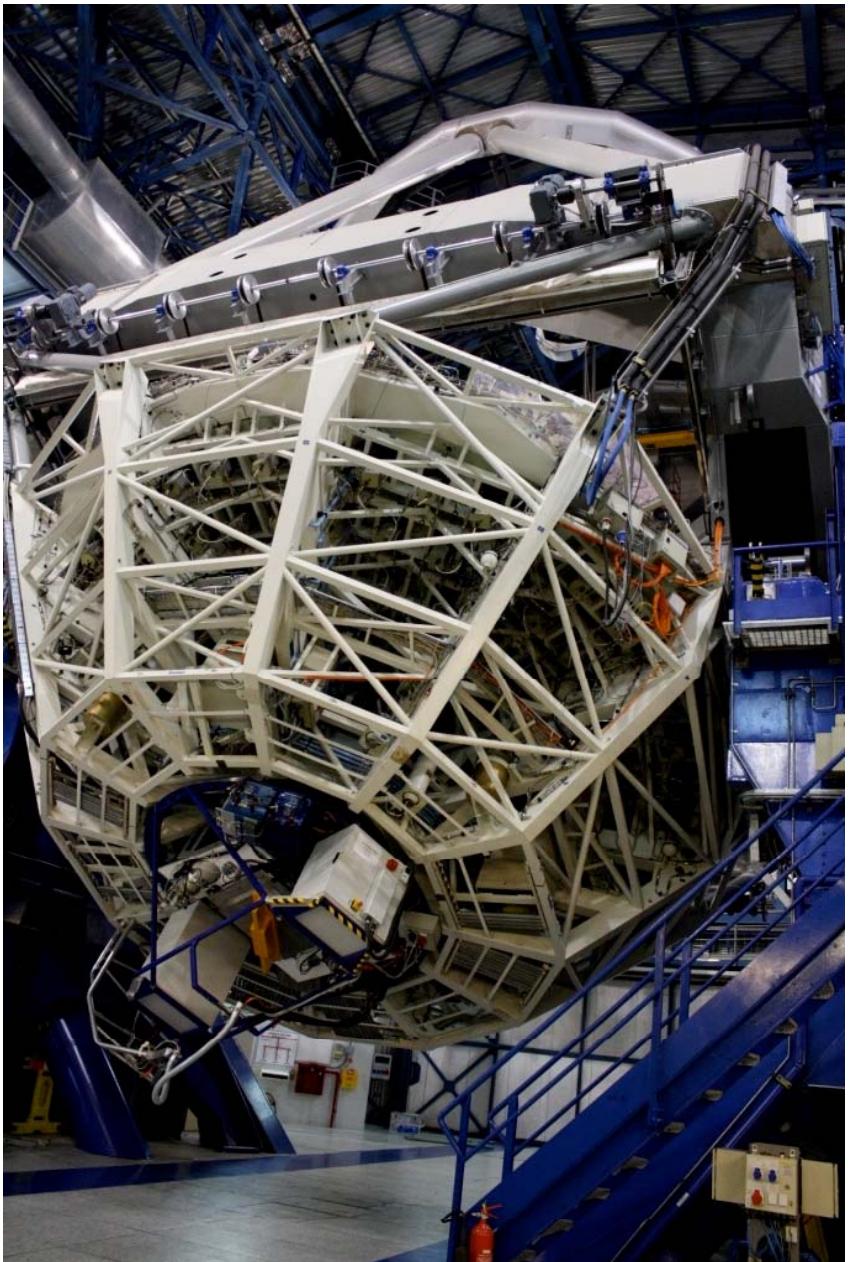


Eventually ...



UT3 integration

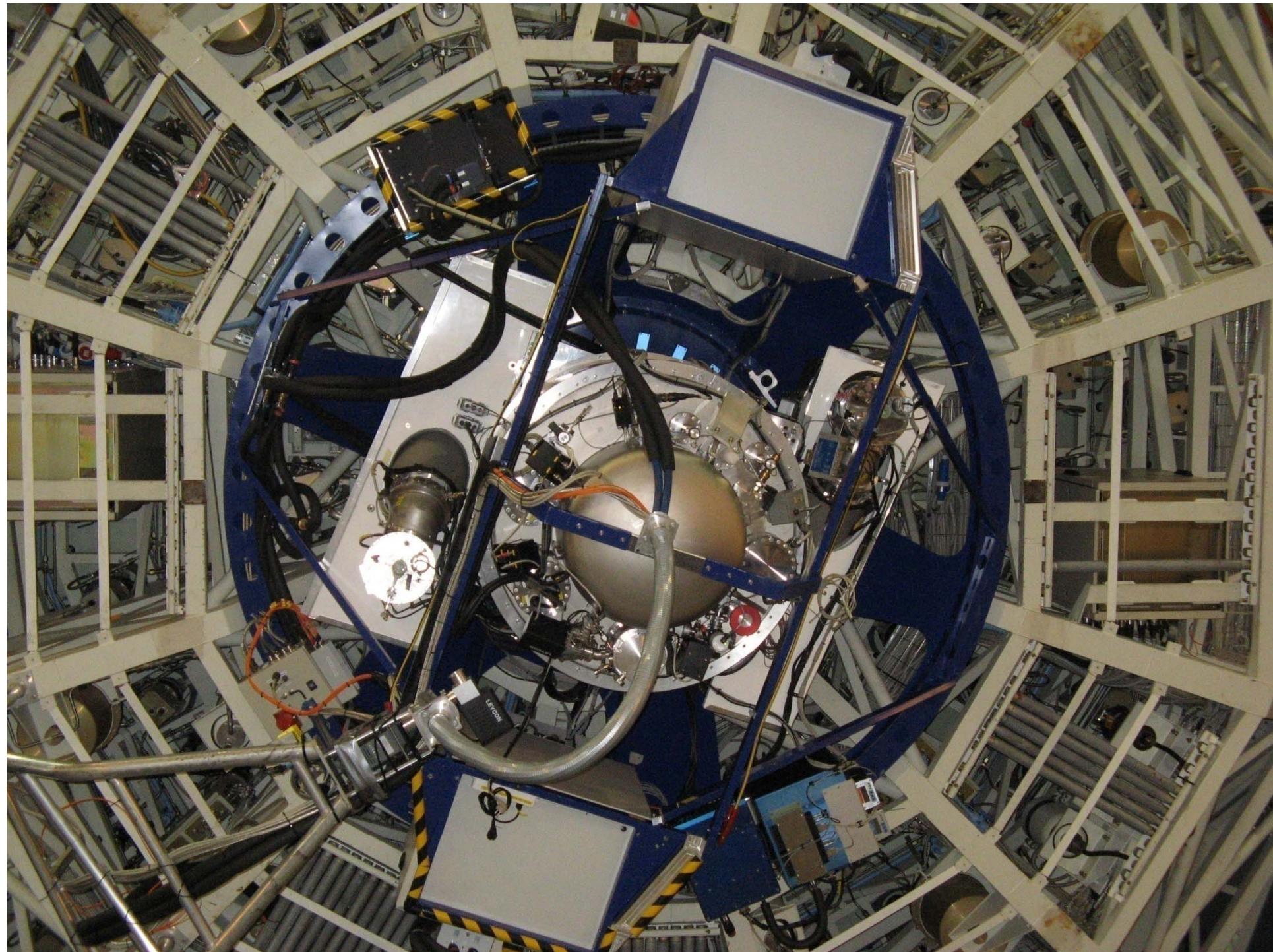




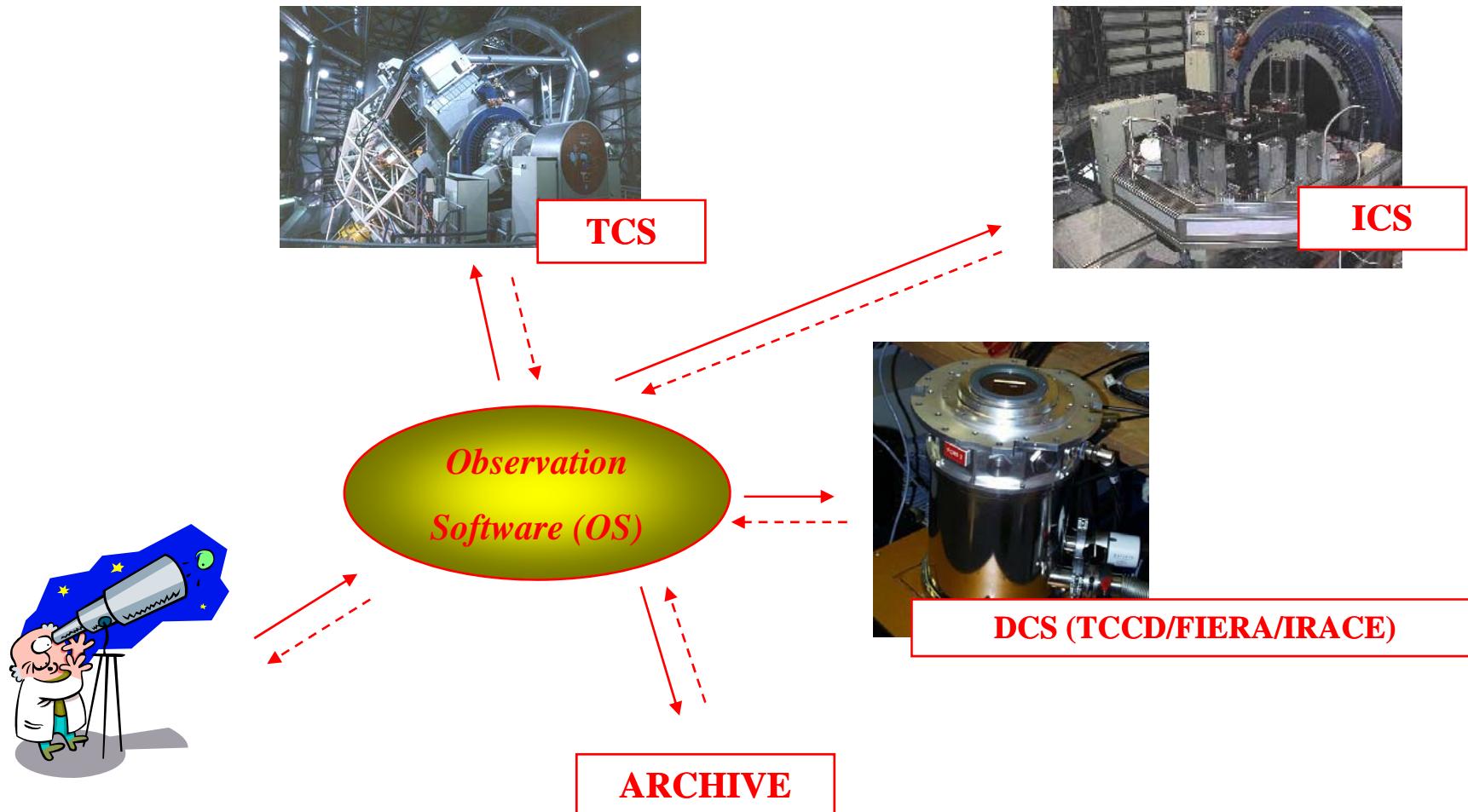
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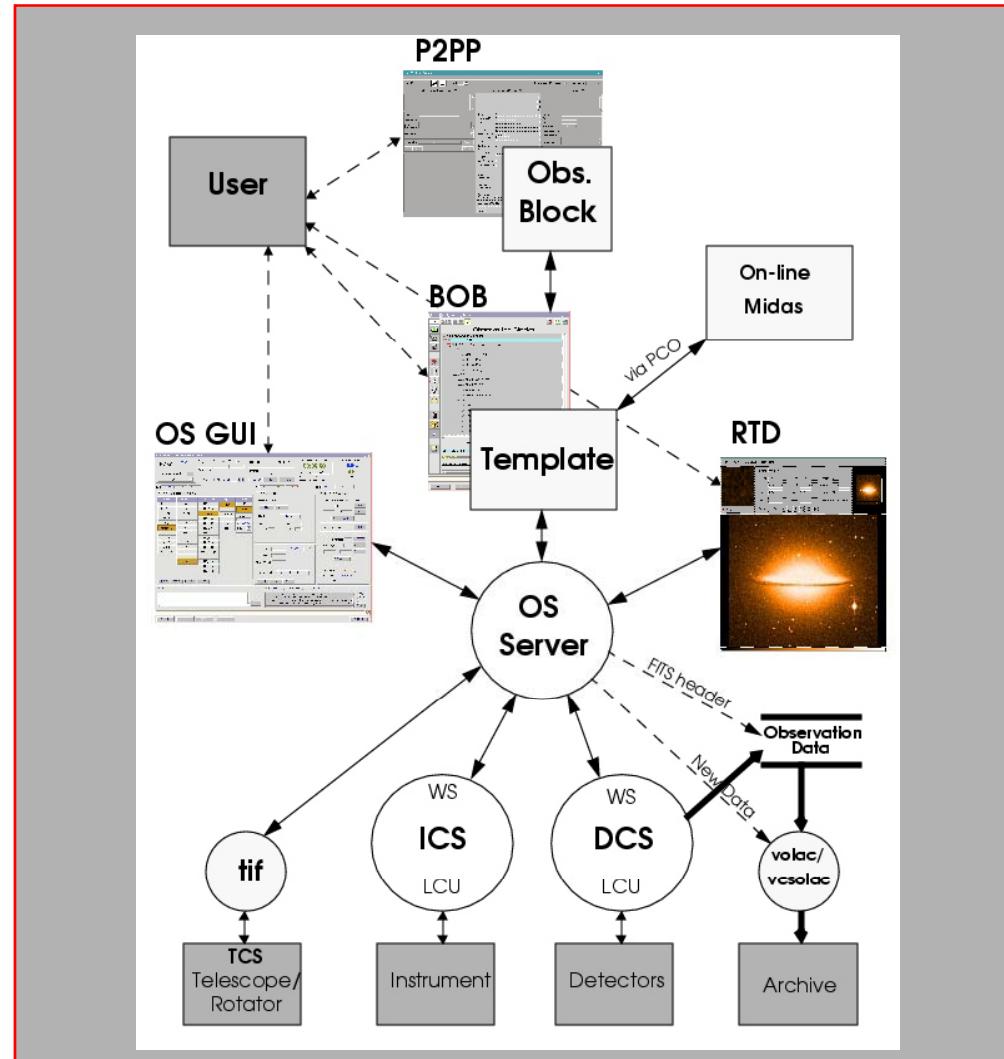
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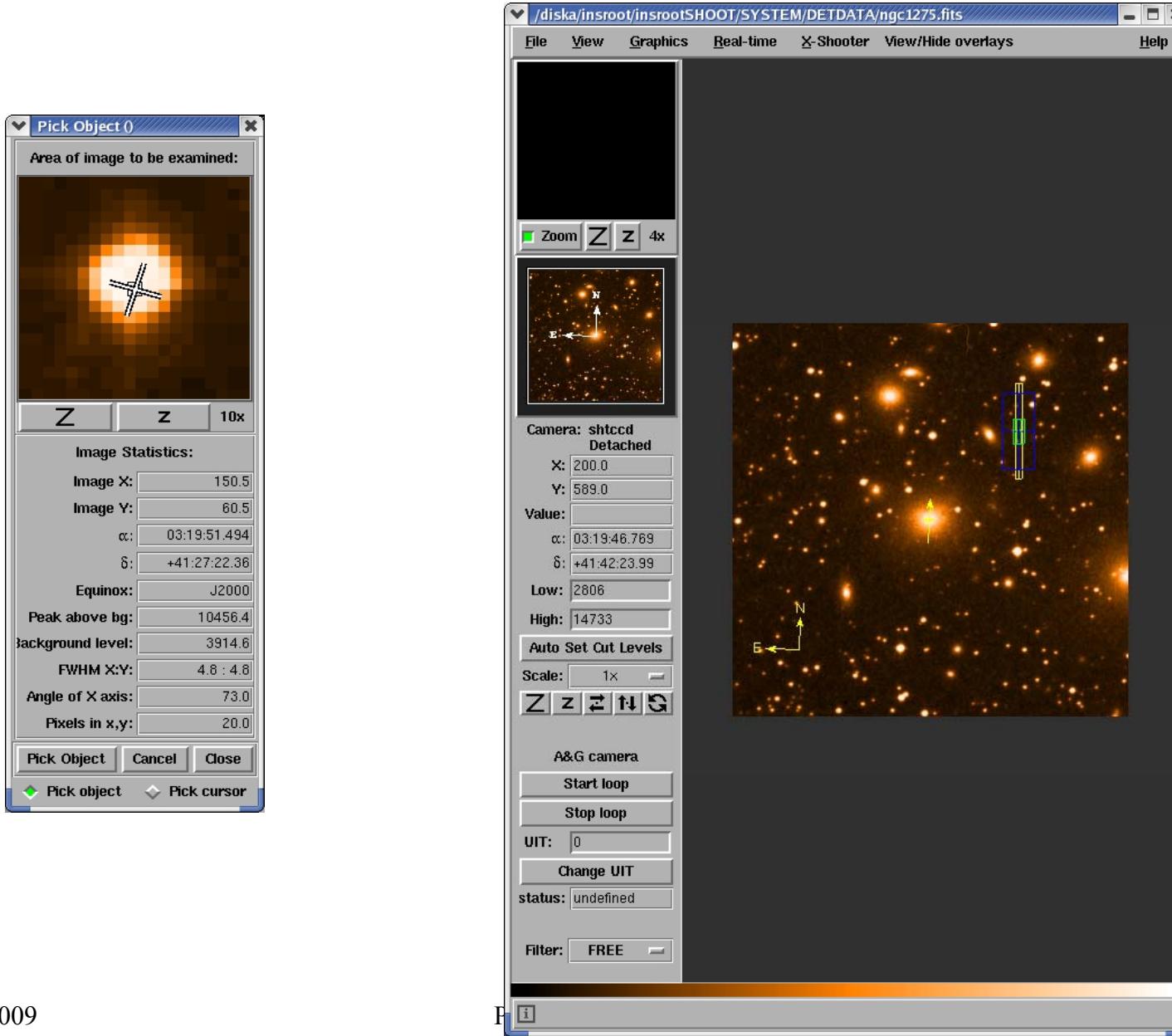
The ESO Control Software



ESO control software - full path



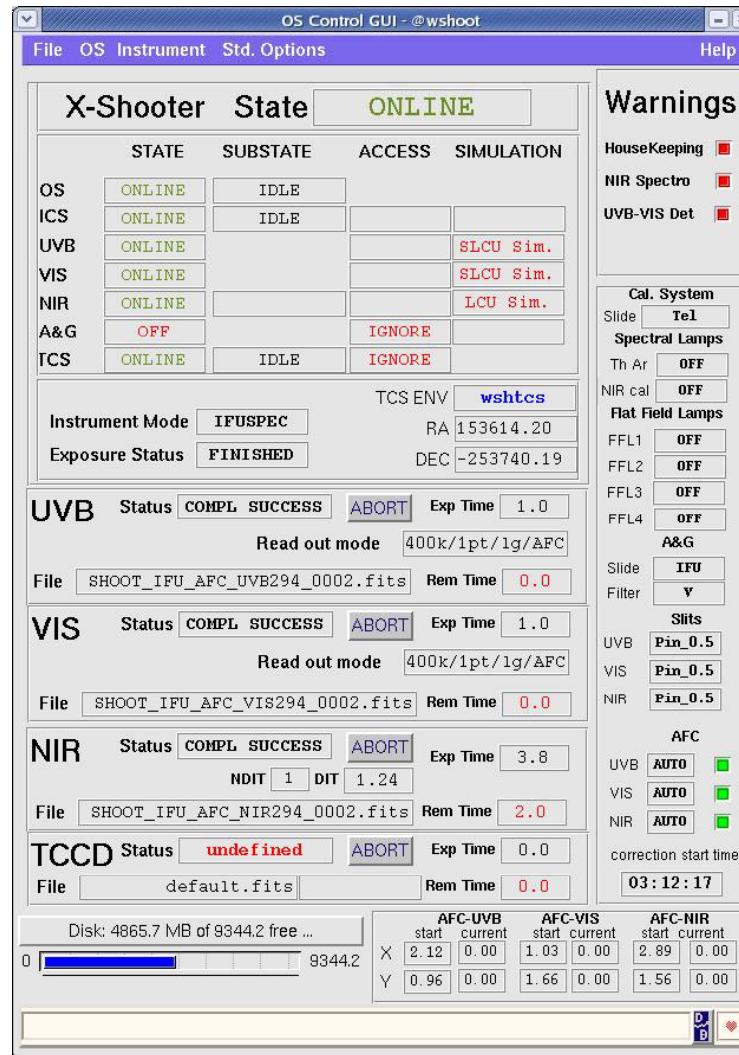
Object Acquisition

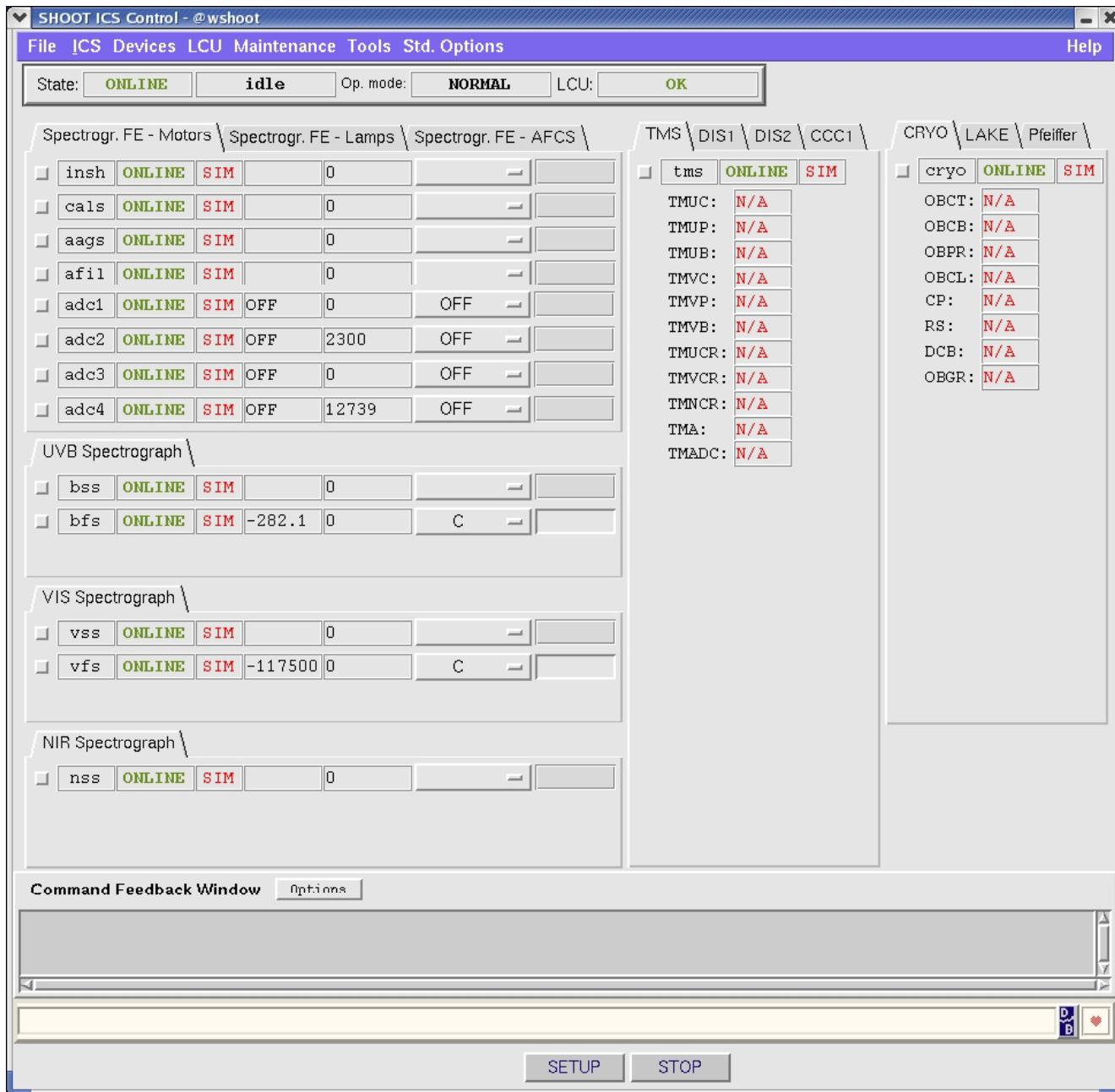




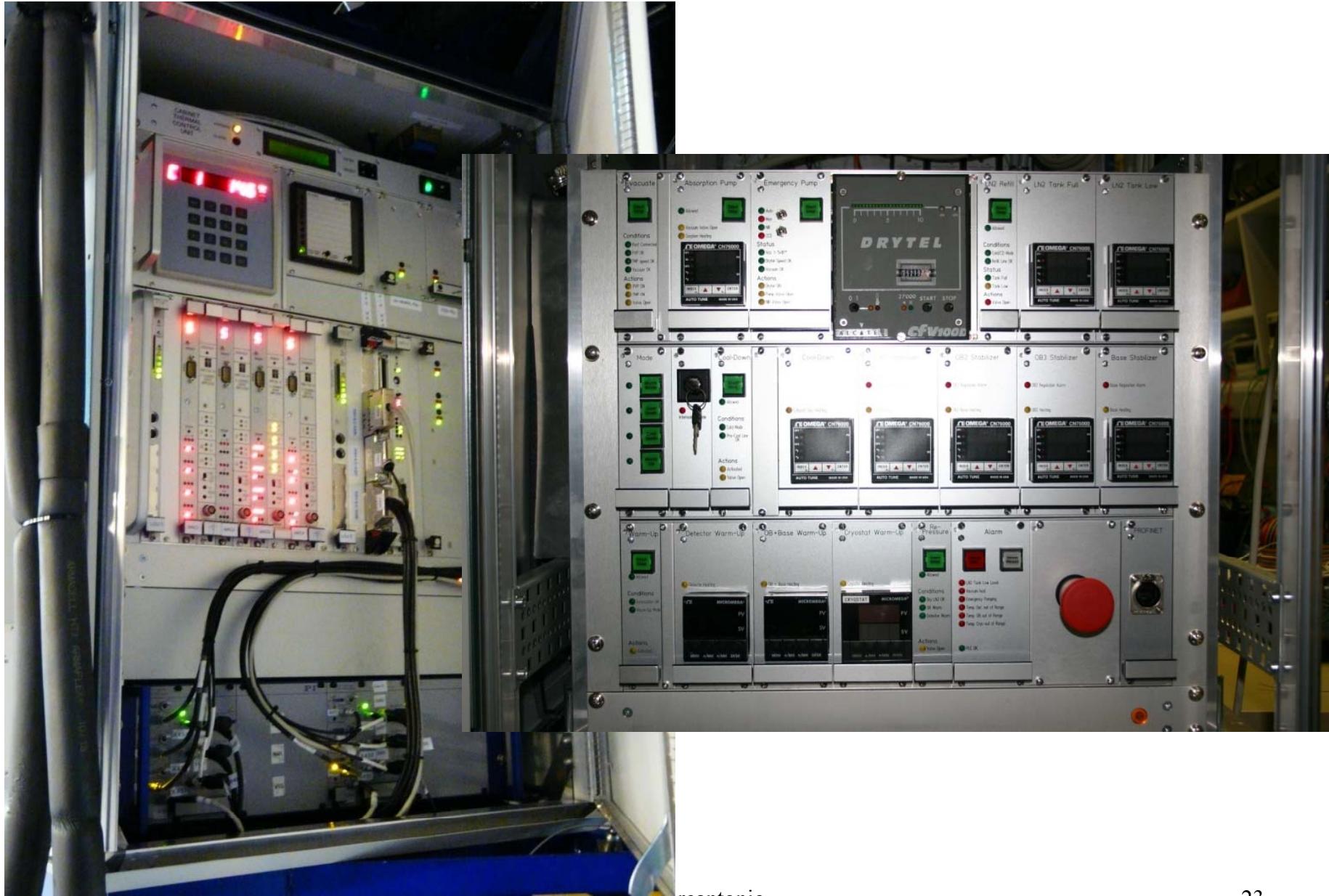
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Observation software (OS)





Local Control Unit (LCU)



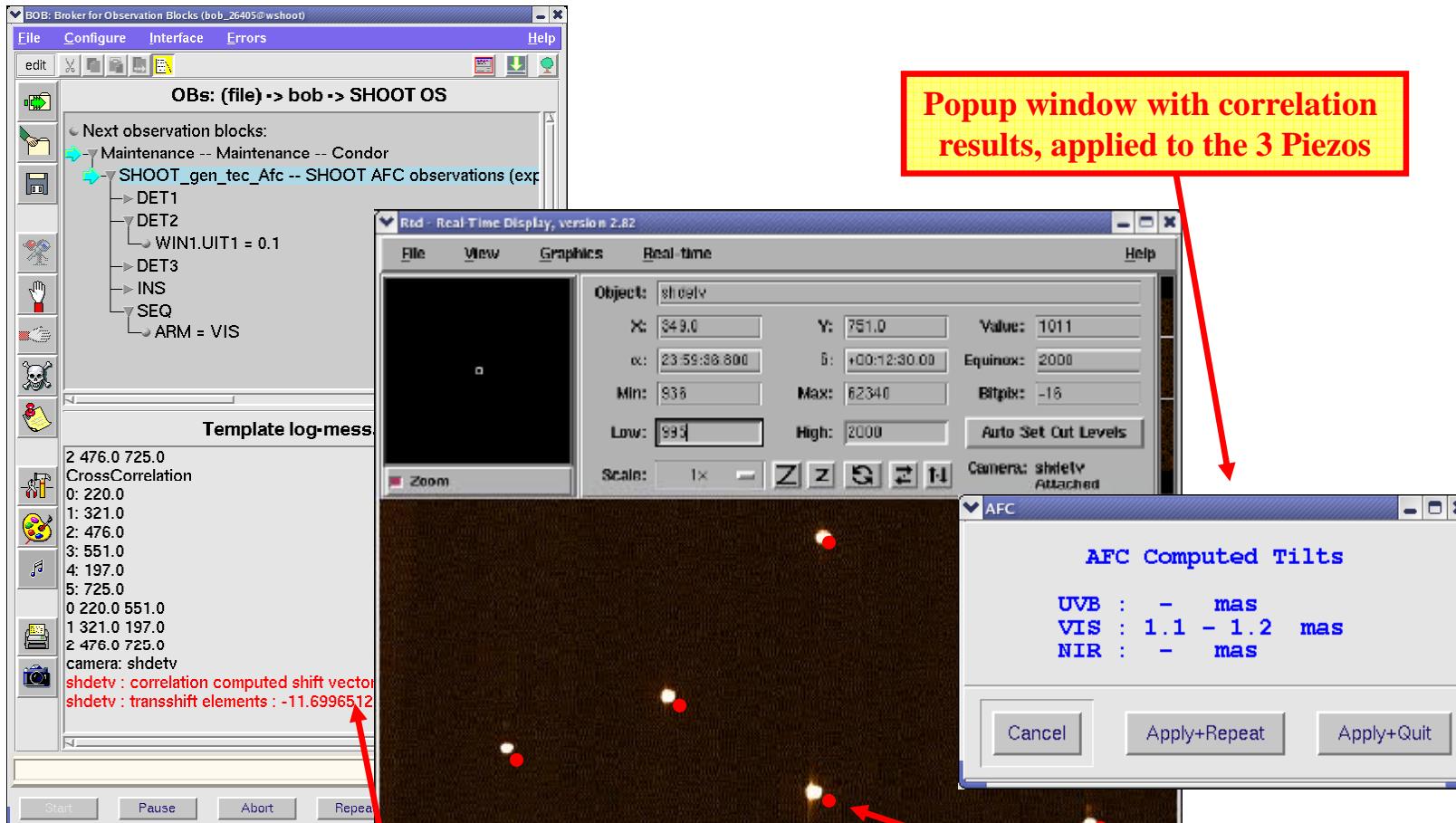
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Active Flexure Compensation



Why AFC?

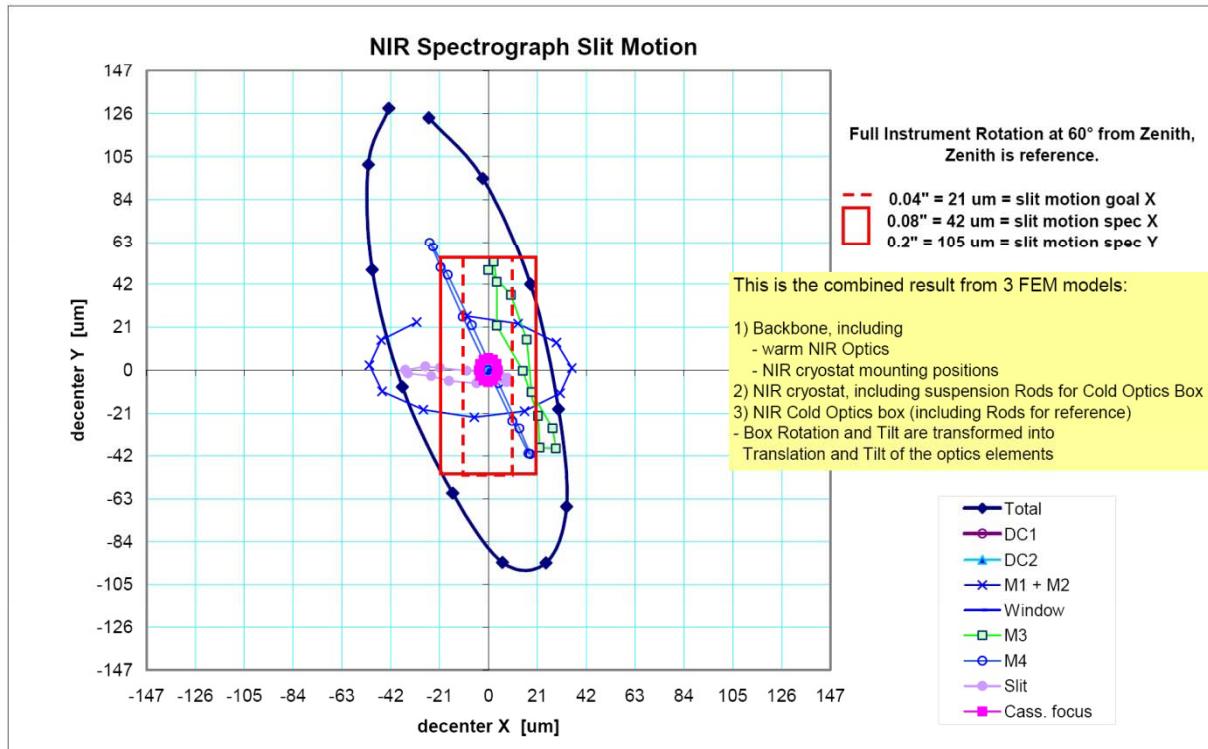
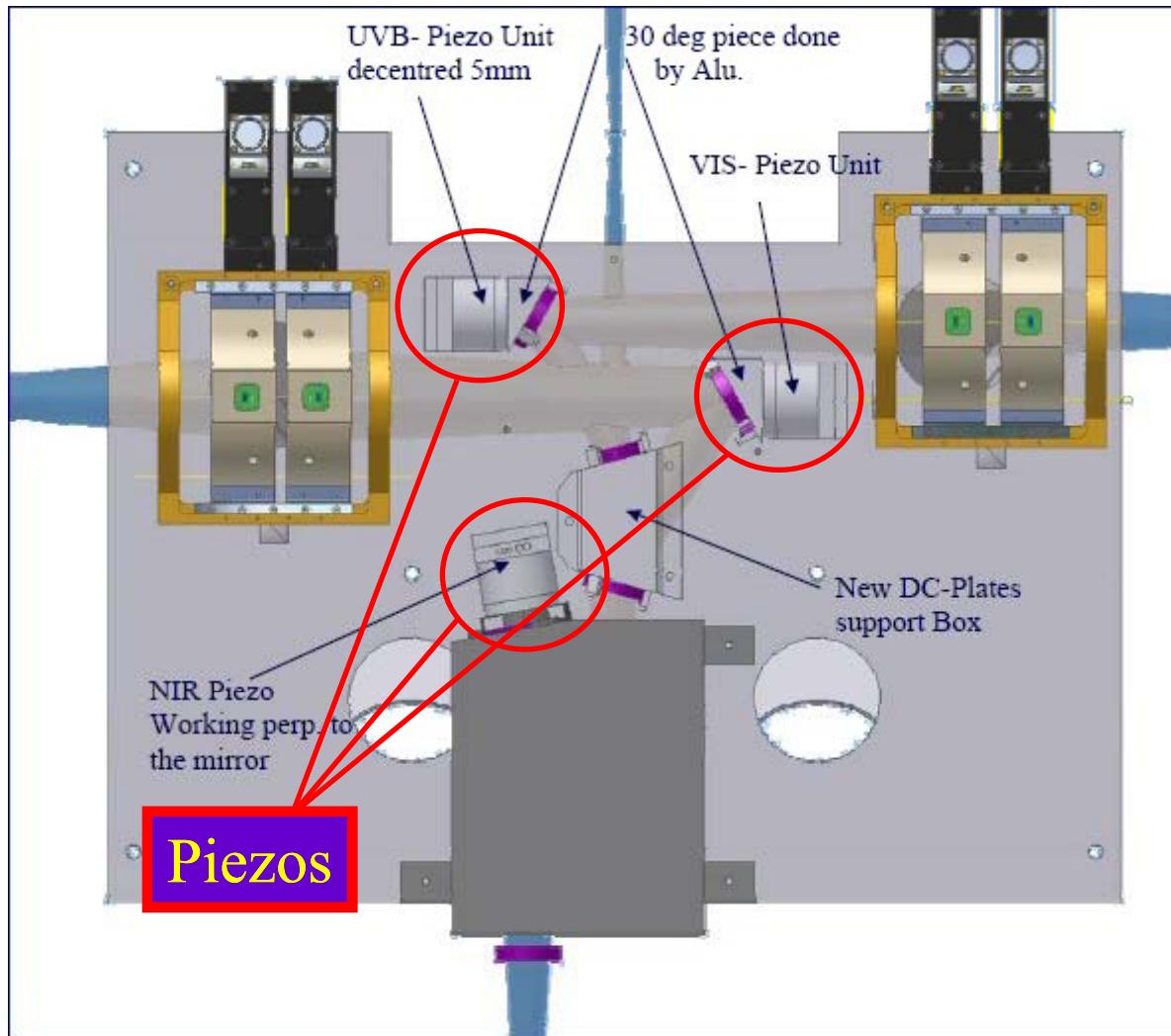
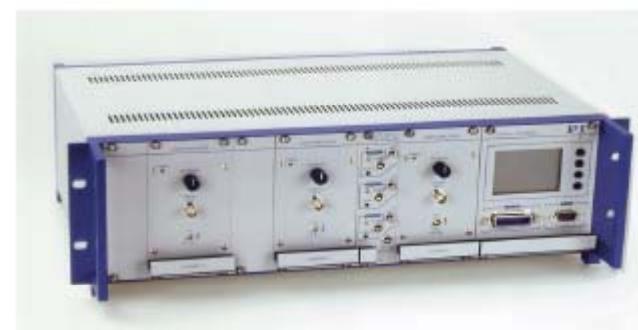
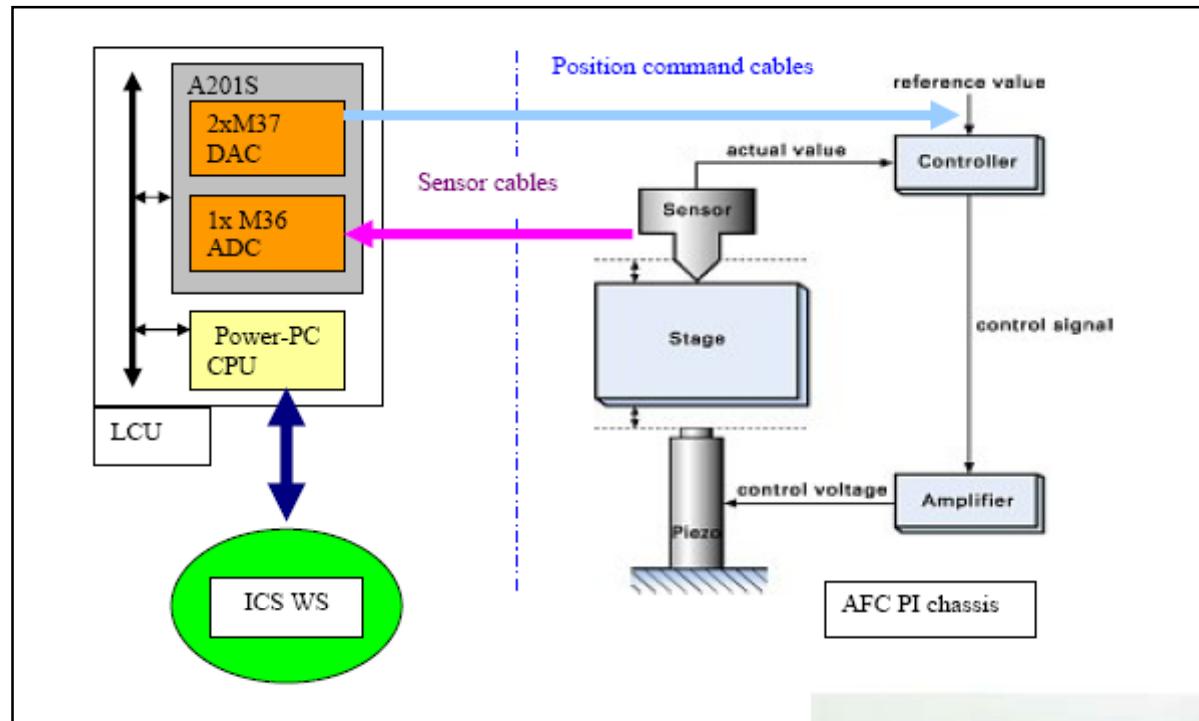


Figure 58 NIR Spectrograph Slit Motion in the NIR slit plane. Combined Result from the 3 FEM analyses.
From Zenith as zero position the maximum NIR slit motion is $-52 \mu\text{m}$ in X direction and $+131 \mu\text{m}$ in Y direction. This is just out of specification.

AFC mechanics



AFC electronics



Commissioning runs

Some results from the three commissioning periods:

- nov 2008 and jan 2008- integration/installation of the UVB and VIS spectrograph with a dummy NIR;

Cited from a report sent to STC:

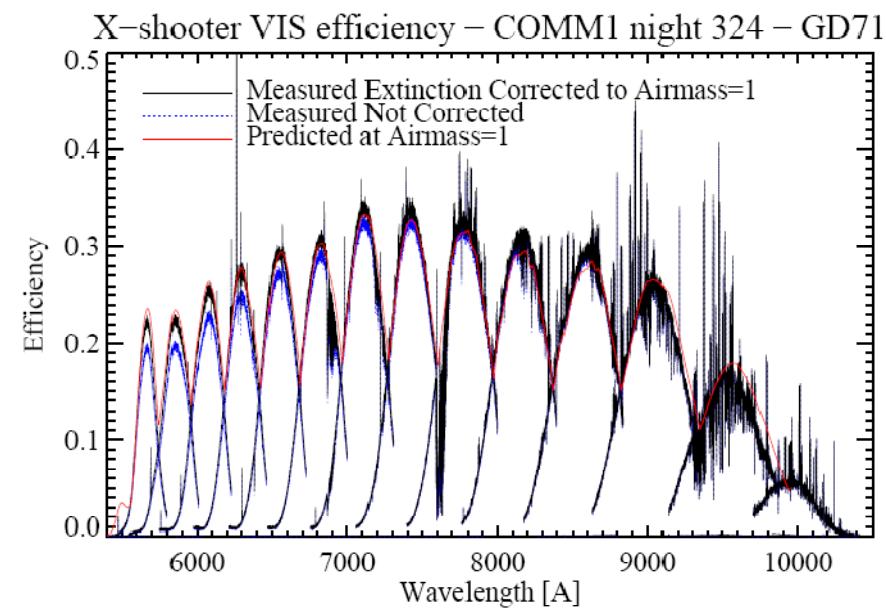
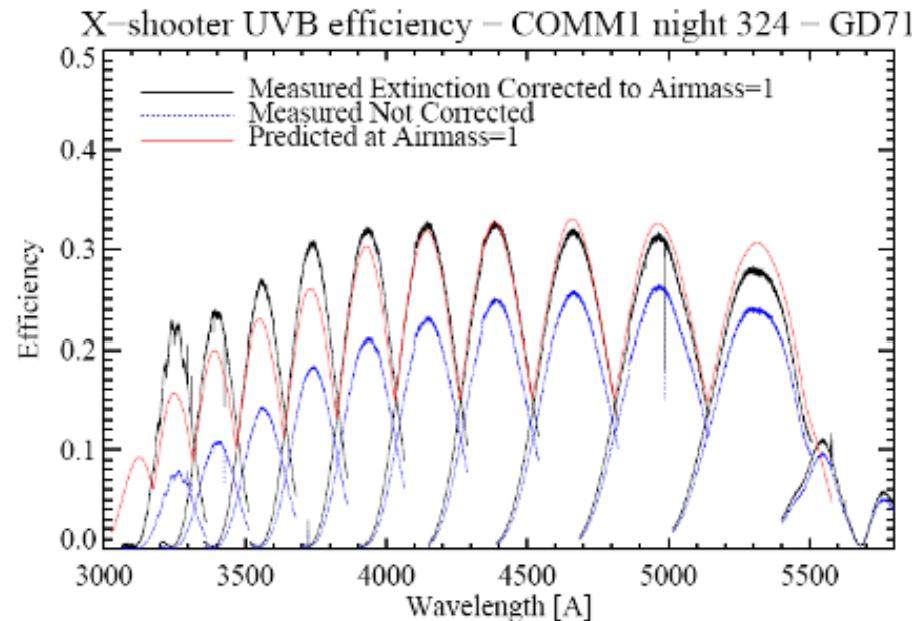
*“... a total of **17 nights**, no losses due to technical problems or weather. The overall efficiency and mechanical flexure are in spec. The hardware, software and operational interfaces with the telescope environment (A&G procedures, OBs, flexure compensation system, ADCs, IFU...) all work as expected in line with the desired “**point and shoot**” objectives for this instrument.”*

- march 2009 – first light with the full instrument:

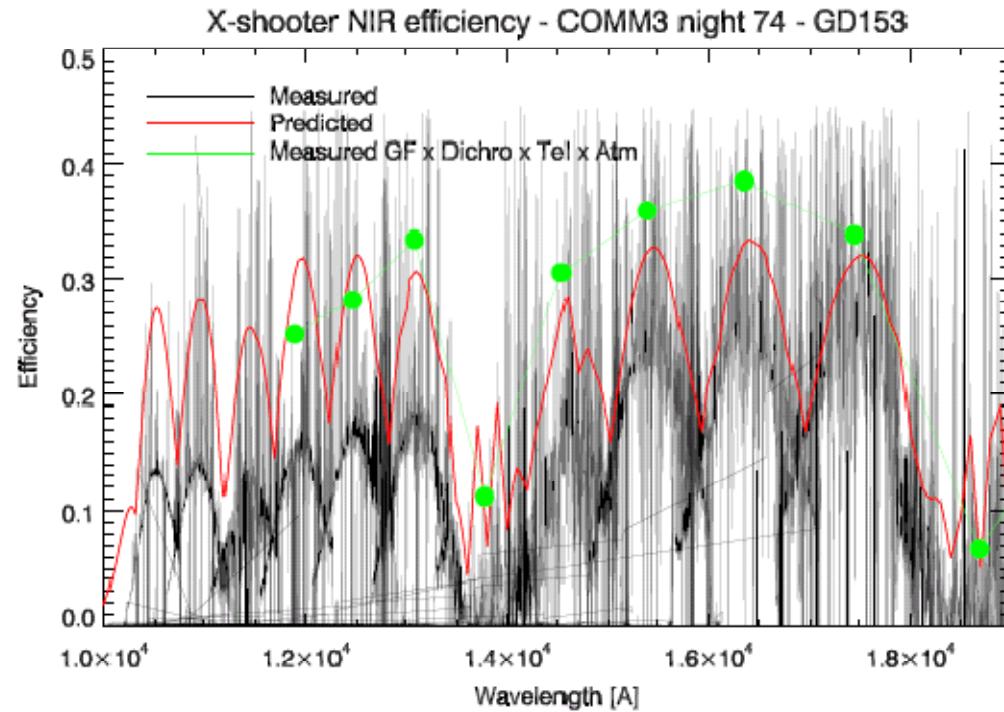
Excerpt from a report sent to STC: **March 14**

*“... **first light was achieved quite smoothly on**. Commissioning continued until the morning of March 20 (**6 nights**, of which 0.75 lost due to VLT SW problems not directly related to the instrument and 1.75 due to weather). There were no significant losses due to problems with specific instrument hard- or software. NIR spectral resolution as a function of slit width was verified and found in agreement with predictions. Nodding and sky offset modes mostly used for the on sky observations. On faint targets, integration times of up to 25 minutes were used, also in the NIR. Classical A – B nodding was tested and works very well in the UVB and VIS arms.”*

Measured Efficiency UVB/VIS

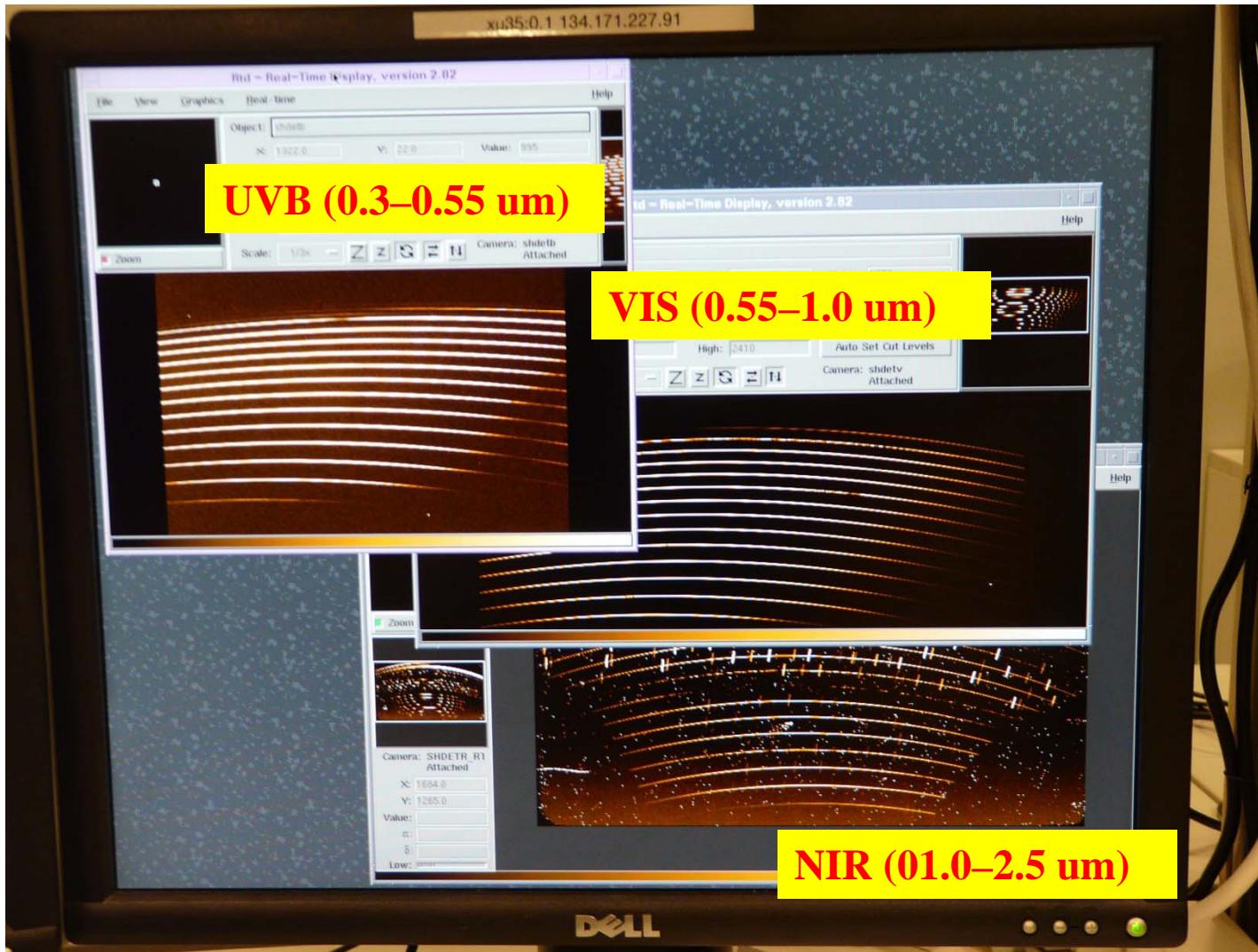


Measured Efficiency NIR



Reduction of the observations of standard stars shows a peak blaze efficiency in U – H band which exceeds or is equal to predictions, except in J band. K band is still being analyzed.

First light 14/03/09 23:22 standard star, 10s

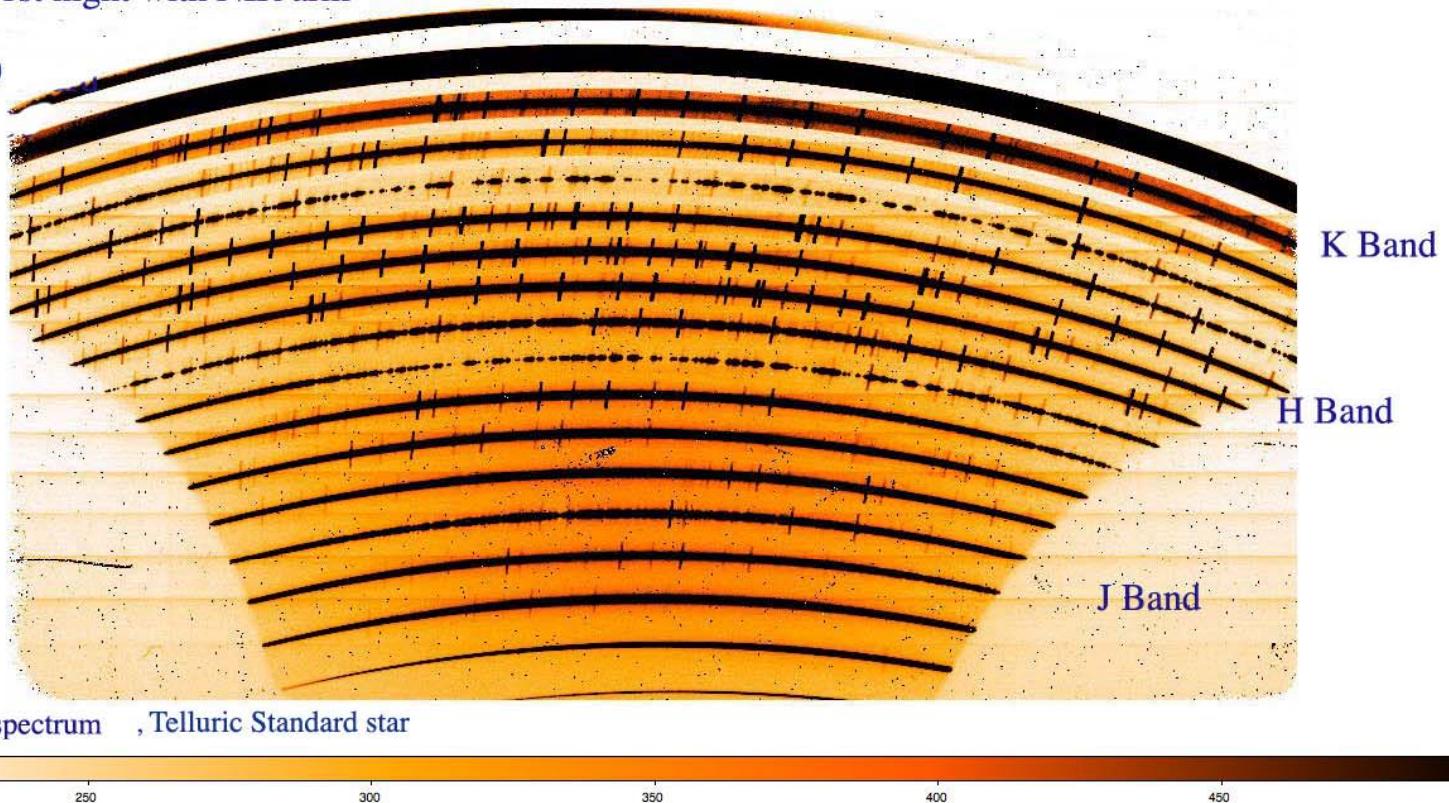


NIR arm J, H, K

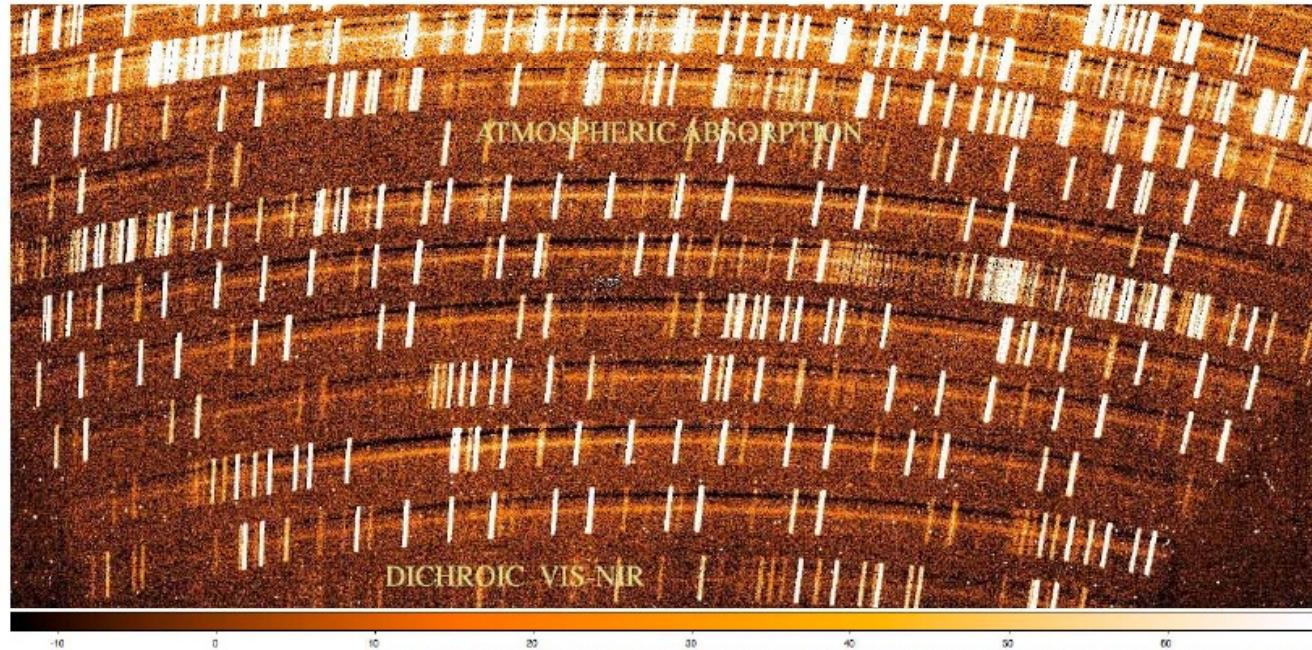
X-shooter 1st night with NIR arm

14.3.2009

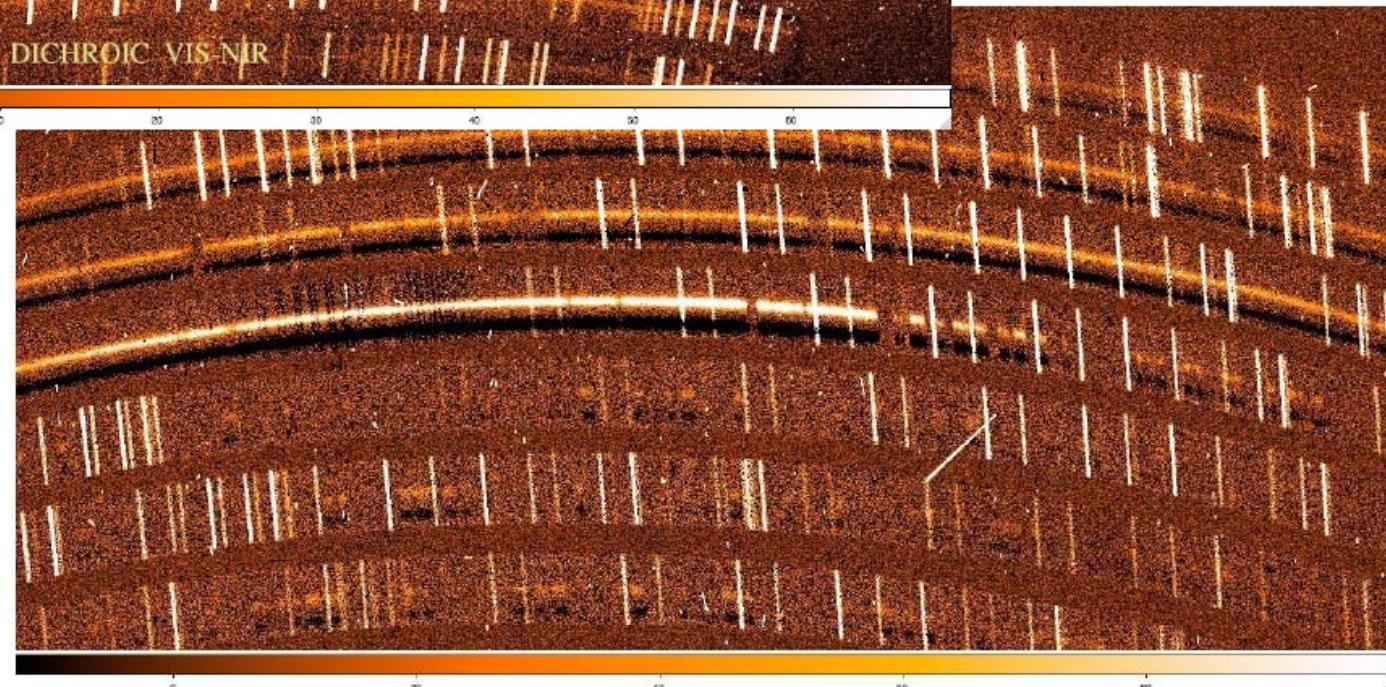
2m, R-5600



Observations of faint targets: a QSO at z=6 (2x30m, subtracted one from the other)



VIS-RED above 700nm,
R= 8800



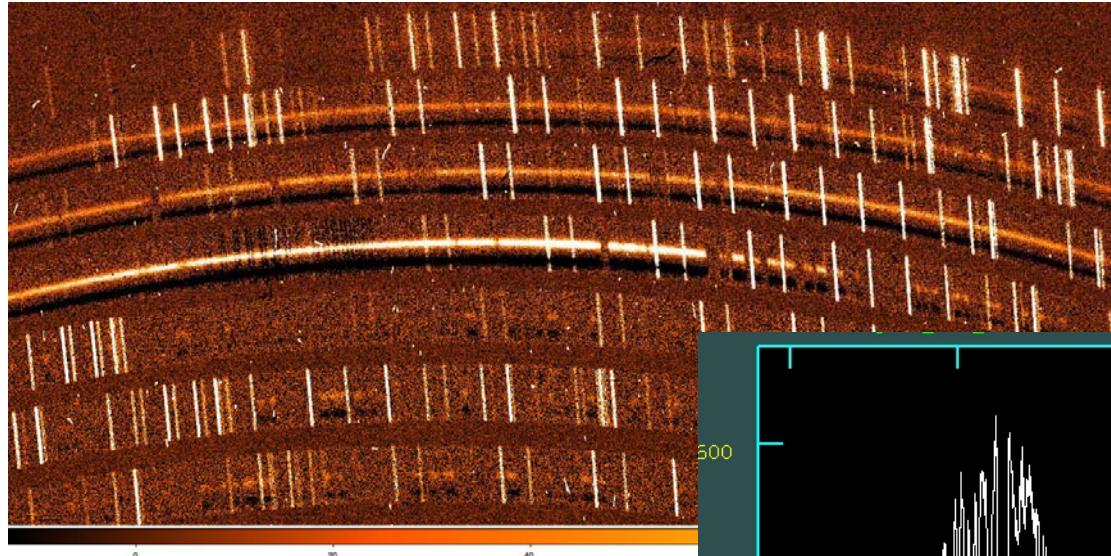
H and J Bands,
R= 5900

N.B. Residual sky due to
intensity variations in the
night



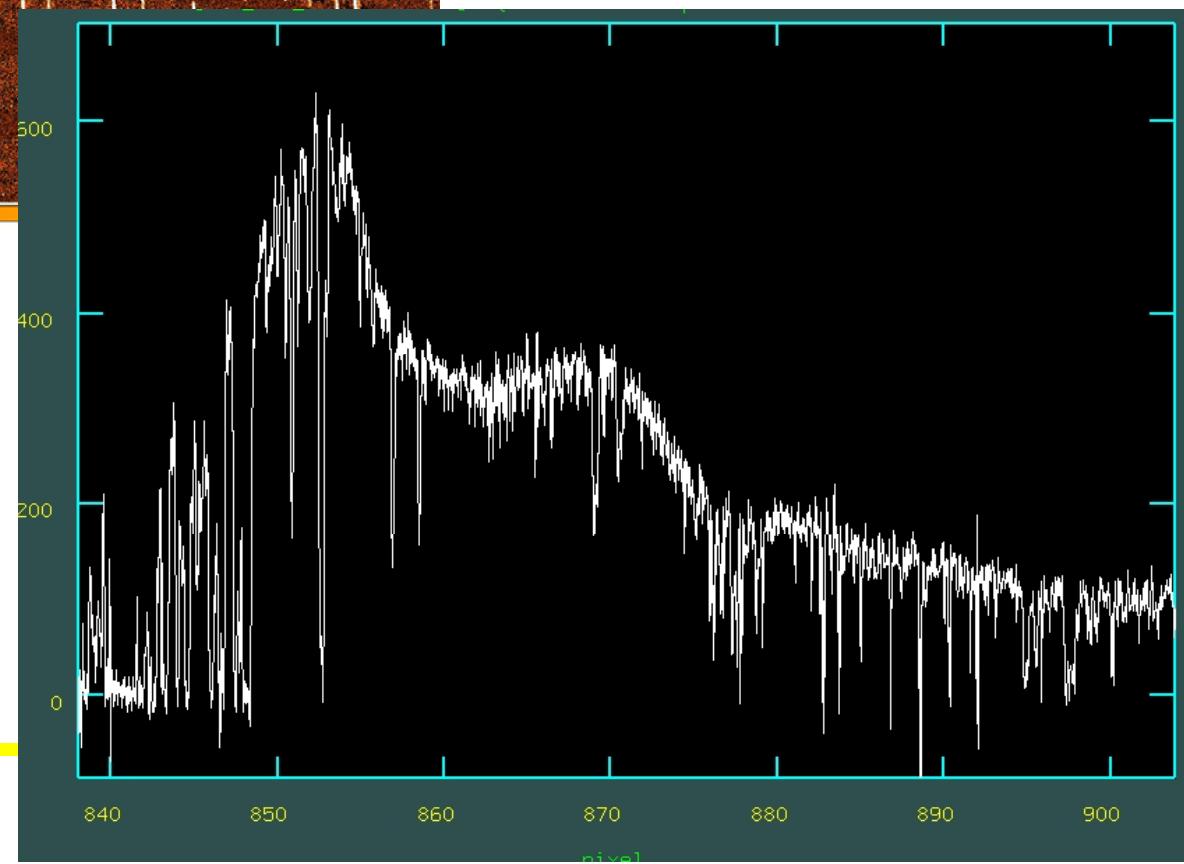
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QSO at $z=6.016$

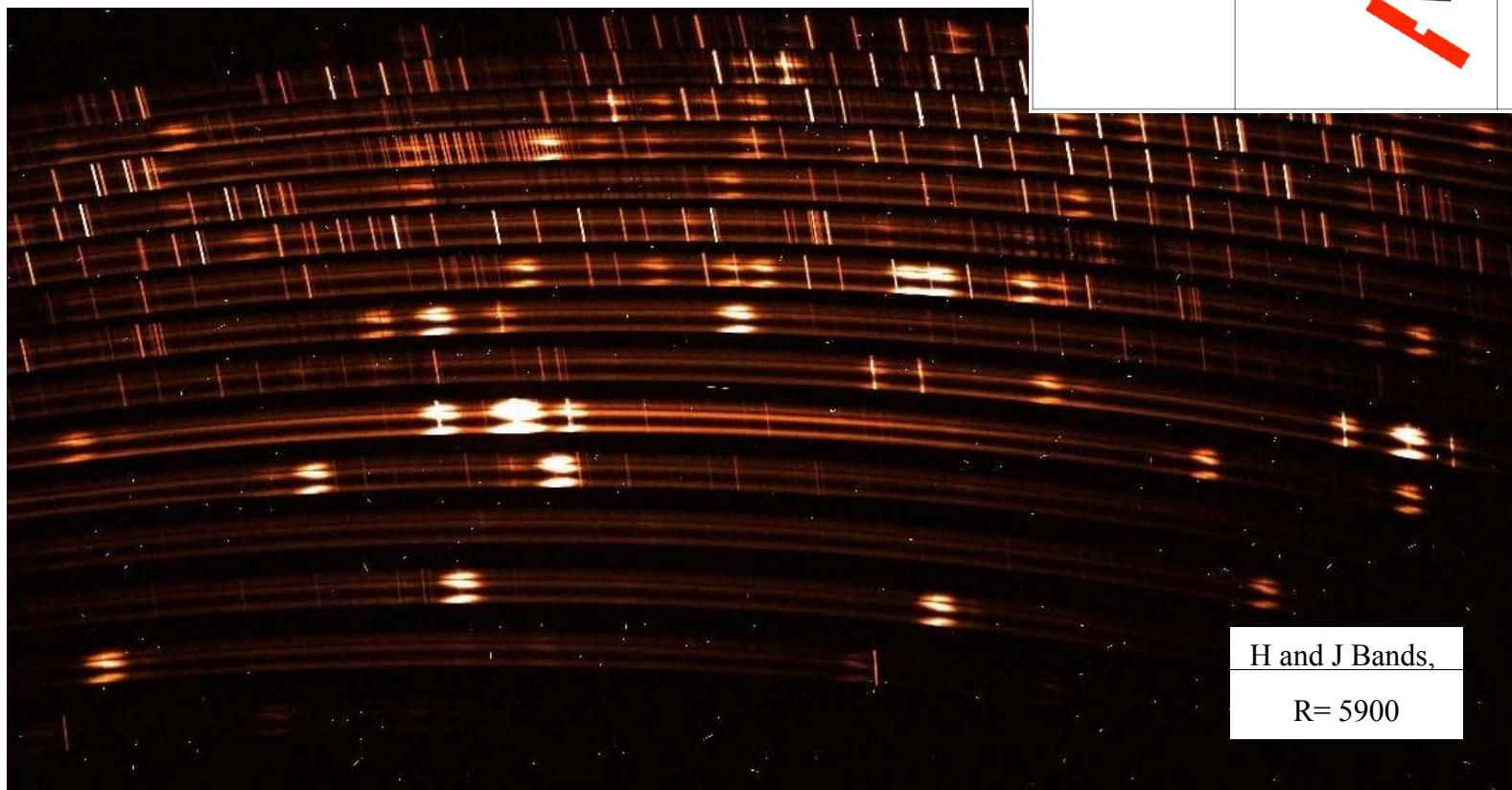


VIS-R above 700nm,
 $R = 8800$

04/05/2009



IFU Vis-R spectrum of SN 1987A in the LMC



20min integration, Spectral Range 550-1025 nm, R=12600

